A novel approach to interaction matrix calibration at THEMIS: accuracy analysis and extension to a temporal model

Clémentine Béchet^{*1}, Michel Tallon², Eric Thiébaut², Isabelle Tallon-Bosc², and Bernard Gelly³

¹Centre de Recherche Astrophysique de Lyon – École Normale Supérieure - Lyon, Université Claude Bernard Lyon 1, Institut National des Sciences de l'Univers, Centre National de la Recherche Scientifique, Institut national des sciences de lÚnivers, Centre National de la Recherche Scientifique, Institut national des sciences de lÚnivers, Centre National de la Recherche

³Télescope héliographique pour l'étude du magnétisme et des instabilites solaires – Centre National de la Recherche Scientifique – France

Abstract

Performance and stability of a closed-loop Adaptive Optics (AO) system are strongly related to

the ability of the controller to model the response of the wavefront sensors to commands sent to the

deformable mirrors. In most AO systems, this is modelled by the so-called Interaction Matrix (IM)

which has to be calibrated.

While most AO systems calibrate the IM on an internal source, implementing the deformable mirror

(DM) in the telescope requires more and more systems calibrate the IM on sky. But this difficulty

brings advantages. For example, the calibration can take into account source specificities such as the

elongation of laser guide stars or the structure of the solar granulation. Furthermore, if the calibration

is done at the system full speed, temporal effects can also be calibrated, such as the temporal response

of the DM.

This is the strategy followed on THEMIS AO, relying on original solutions for its wavefront sensing,

its control, and its calibration, each one based on an inverse problem approach. We present

*Speaker

in this

work the interaction matrix calibration and we analyze the quality and features that can be obtained

from this original approach.

In THEMIS AO, the IM is calibrated during open-loop observations of the sun granulation using a

binomial random distribution of perturbation commands. The properties and benefits of this approach

are analyzed. The uncertainties on the obtained IMs perfectly follow the theory. Data are compared

to theory in Fig. 1 (left) showing noise decrease as ($\epsilon_{\rm v}/{\rm nt}){-1}$ with nt the number of measurements and

 ϵ the ratio between the standard deviations of the perturbation and of the turbulence.

For THEMIS AO, residual calibration errors suggest that a time dependent model of the IM is

needed. To account for the temporal response of the DM in the controller, our calibration approach

has thus been extended to fit an IM model consisting in a sequence of matrices. We show that, in

practice, 2 components are sufficient and that the second IM matrix is well approximated by the first

one times a simple factor. Figure 1 (right) shows the relative contribution of the 2-component IM

fitted on THEMIS calibration data. Our time dependent IM model thus amounts to applying the

static IM matrix to the current commands plus a fraction (11% here) of the previous commands.

THEMIS calibration approaches and results are of importance for other high-frequency AO systems,

particularly for High-Contrast AO systems.

Keywords: calibration, on, sky, optimisation, temporal interaction matrix